## Modern Mathematical Methods in Relativity Theory II

## Tentative Table of Contents [and Reading List]

- 1.-2. World lines of extremal length as geodesics [MTW 13.4].
  - 3. The equation for a geodesic: bridge between physics and geometry. Inertial force  $\neq 0$  $\iff$  coordinate frame is curvilinear; equivalence principle: gravitational force = inertial force. "geometry" = "gravitation".
  - 4. Momentum and energy: its definition and conservation [TW 7.1-7.7].
  - 5. The particle density 3-form [MTW Box 4.4, Box 5.2, Box 15.1.f]; particle 4-current: density and flux.
  - 6. Stress-energy tensor: the flow of momenergy; the physical significance of its components [MTW Ch.5].
  - 7. Conservation of momenergy; creation of momenergy: (a) expressed in physical terms; (b) expressed as an integral over the boundary [MTW Box 15.1.B] of a 4-volume.
  - 8. Creation of momenergy: (c) expressed as a 4-volume integral: Gauss's theorem for a vector valued integral; (d) expressed as the exterior derivative of a vector valued three form.
  - Conservation laws formulated in terms of the generalized Stokes theorem [MTW 5.8, Box 5.3, Box 5.4]; Summary: Conservation of momenergy expressed at four levels of mathematical generalization [MTW Box 15.2]; Expansion of a moving volume of fluid [MTW 22.1-22.3].
  - 10. Perfect fluid and its equation of motion as implied by momenergy conservation. Energy conservation, particle conservation, and the chemical potential. Various manifestations of  $T^{\nu}_{\mu;\nu} = 0$  [MTW 22.-22.3].
  - 11.  $\partial \partial V = 0$  and the Einstein field equations [MTW Ch. 15]; examples of  $\partial \partial V = 0$ : div curl=0, Bianchi identities.
  - 12. Vectorial form of Stokes' theorem: the 1-2 version. Jacobi's identity [MTW Ex. 9.12 a and c] and the infinitesimal Gauss's theorem revealed by a chipped cube.
  - 13. Gauss's theorem as a bridge from  $\partial \partial V = 0$  to the Bianchi identities.
  - 14. Moment of rotation per volume= Einstein tensor [MTW 15.4].
  - 15. Moment of rotation, moment of force, and the Einstein field equations.
  - 16. Einstein's equations  $\Rightarrow$  conservation of momenergy [MTW ch. 15]; integral form of the Einstein field equations; comparison with integral formulation of Coulomb's law and Ampere's law. Spherically symmetric systems.
  - 17. Einstein's equations for spherically symmetric configurations.
  - 18. Geometrical and matter degrees of freedom.
  - 19. Helmholtz's theorem.

- 20. Integration of Einstein's field equation via mass-energy conservation; Mass distribution determines spatial geometry. Inner geometry via imbedded surface; application to the space geometry of a spherical star [MTW 23.8].
- 21. Simplified Einstein field equations. Equations of hydrostatic equilibrium [MTW ch.23]; equilibrium configurations: stable vs. unstable [MTW ch. 24].
- 22. Hamilton-Jacobi theory and the principle of constructive interference [MTW Box 25.3]. Constructive interference  $\Rightarrow$  world lines have a finite length determined by Planck's constant. Derivation of Heisenberg's indeterminacy principle.
- 23. Reconstruction of classical world lines from the principle of constructive interference.
- 24. Hamilton- Jacobi analysis of the orbits of a particle in the spacetime of a spherically symmetric vacuum configuration [MTW 25.5, Box 25.4].
- 25. Precession of the perihelion and the deflection of light by the sun [MTW 25.5,25.6].
- 26. Schwarzschild spacetime: Regular behavior of proper time, proper distance, and curvature at the Schwarzschild radius r=2M [MTW 31.2]; geometry and topology of two asymptotically flat connected regions [MTW 31.6, 31.7].
- 27. Schwarzschild spacetime: dynamics, causal structure near r=2M; Eddington-Finkelstein coordinates [MTW 31.4, Box 31.2]; Kruskal-Szekers coordinates [MTW 31.5].
- 28. Globally defined coordinate system for Schwarzschild spacetime.
- 29. Scalar, vector, and tensor harmonics, their behavior under parity transformation; geometrical objects on 2-D Lorentz spacetime.
- 30. Spherical tensor harmonics; representation of generic perturbations as "odd" and "even" geometrical objects on  $M^2$ ; Coordinate induced ("gauge") perturbations of a tensor field; gauge invariant geometrical objects.

## **Resource Texts**

- C.W. Misner, K.S. Thorne, and J.A. Wheeler, *Gravitation*, (Freeman, San Francisco, 1973).
- 2. E.F. Taylor and J.A. Wheeler, *Spacetime Physics*., SECOND EDITION (1st Edition)